

National Aeronautics and Space Administration



# Optical Observations of Space Debris

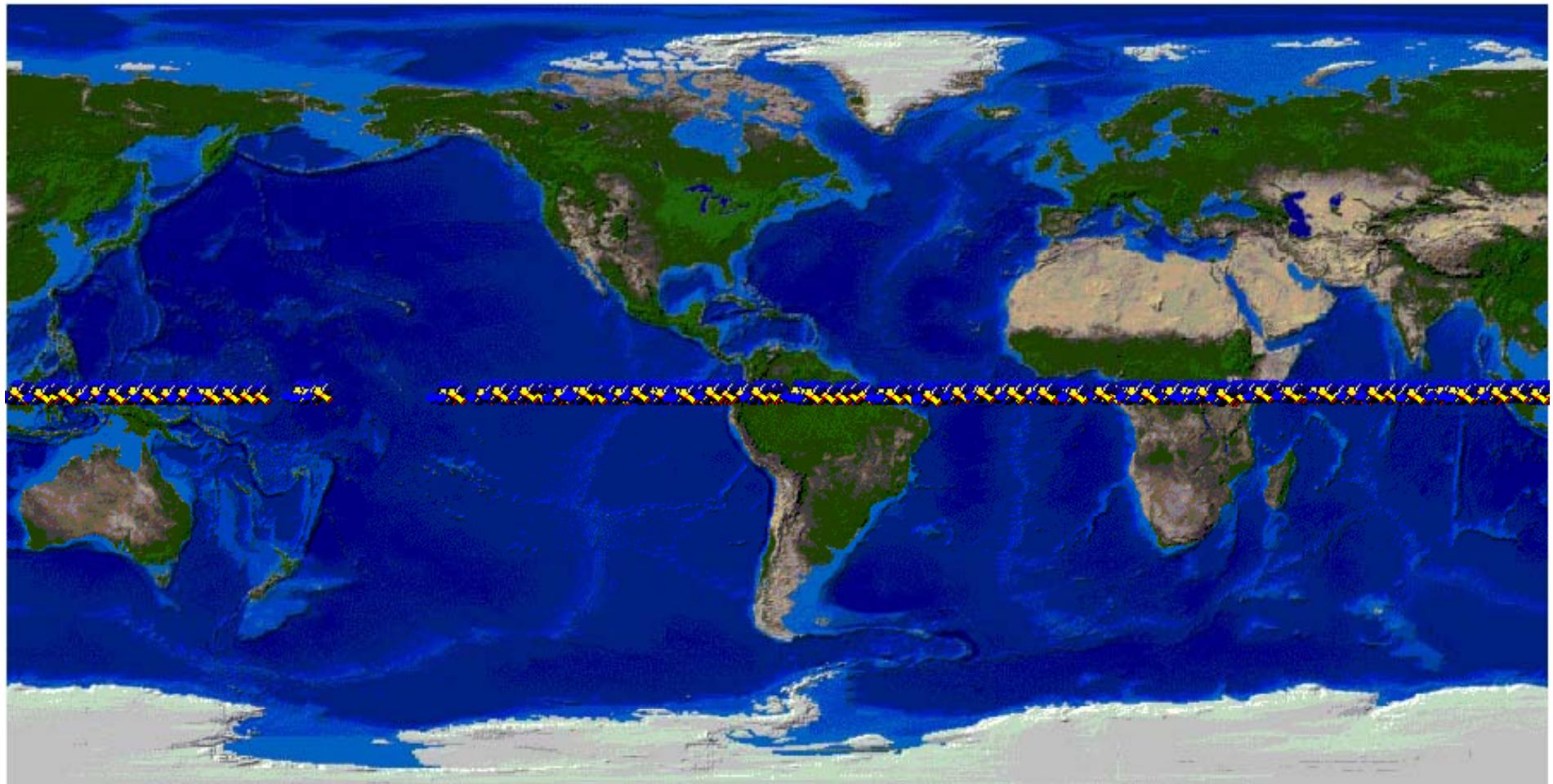
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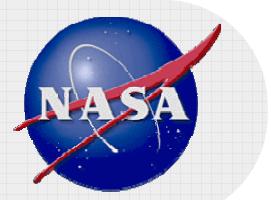
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& Thomas Kelecy (Boeing)

*Supported by*  
*NASA's Orbital Debris Program Office*  
*Johnson Space Center, Houston, Texas*

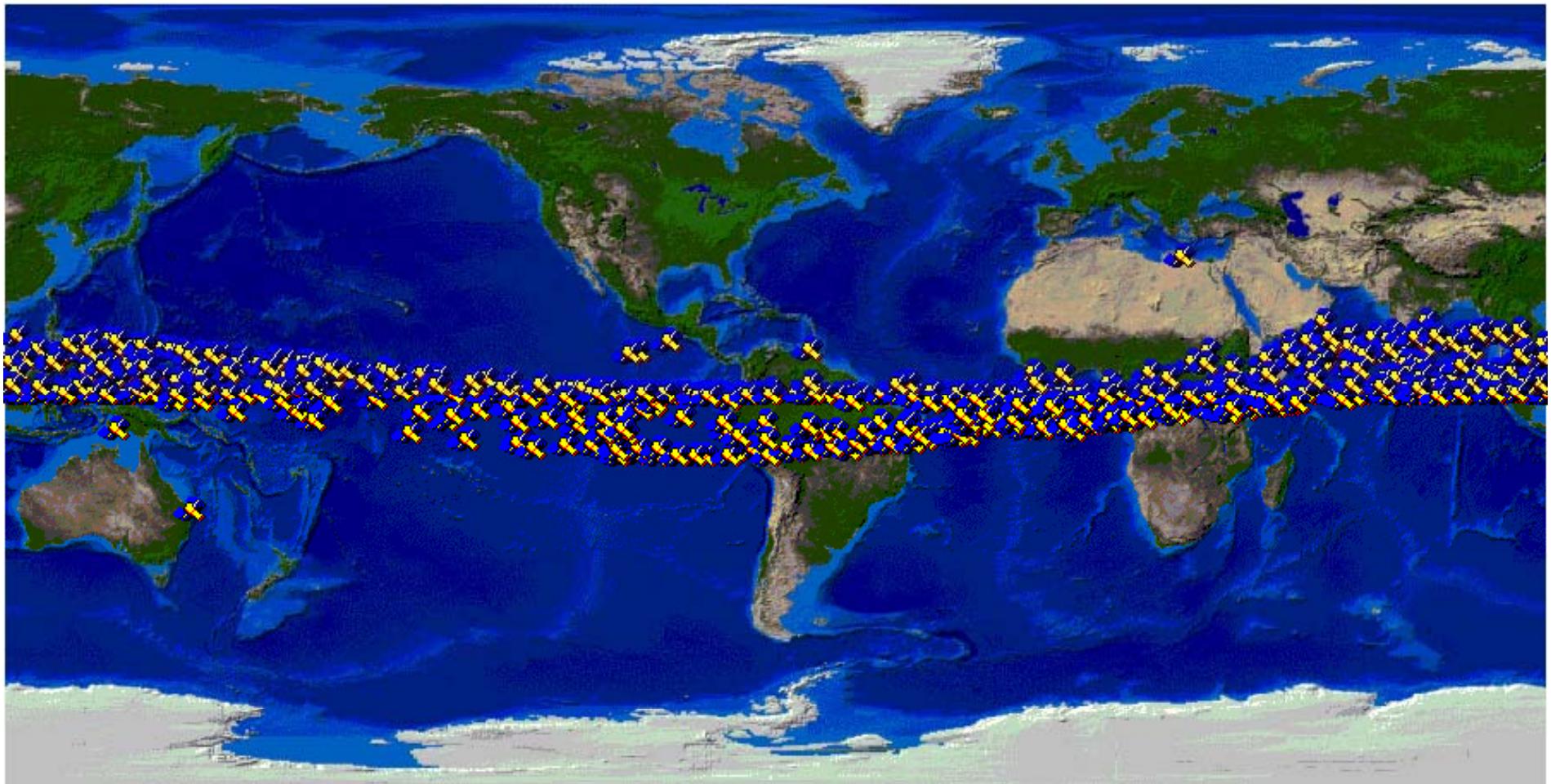


# GEO Stationkeeping Satellites





## All Cataloged Objects at GEO





# Cerro Tololo Inter-American Observatory



**MODEST – Michigan Orbital DEbris Survey Telescope**  
*the telescope formerly known as the Curtis-Schmidt*

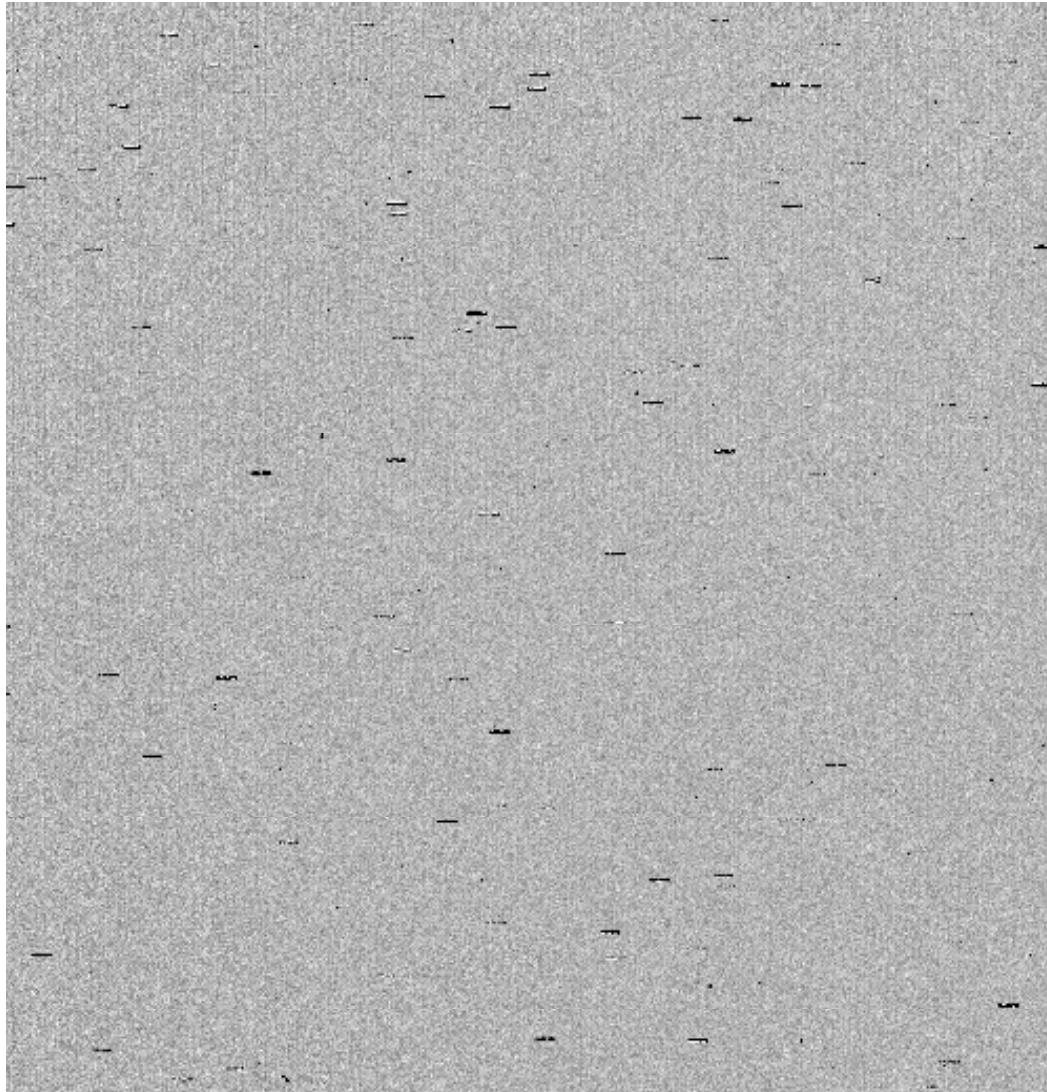


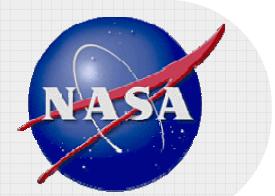
0.61/0.91-m Schmidt telescope  
GEO debris survey began February 2001



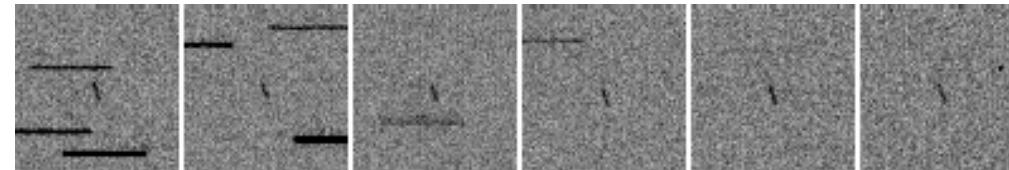
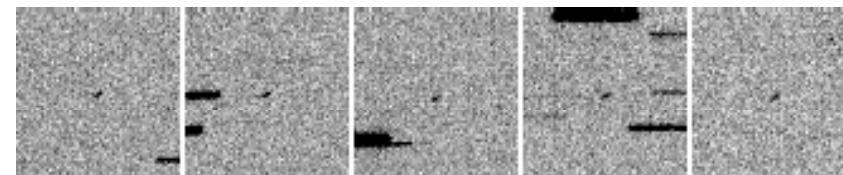
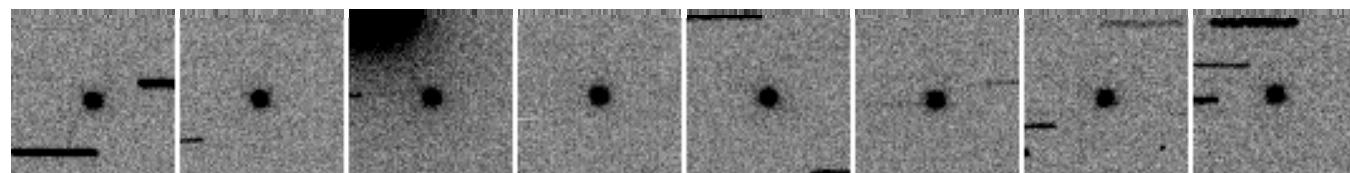


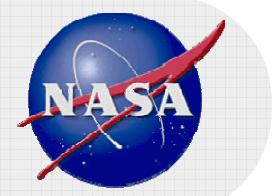
## MODEST Data Sample



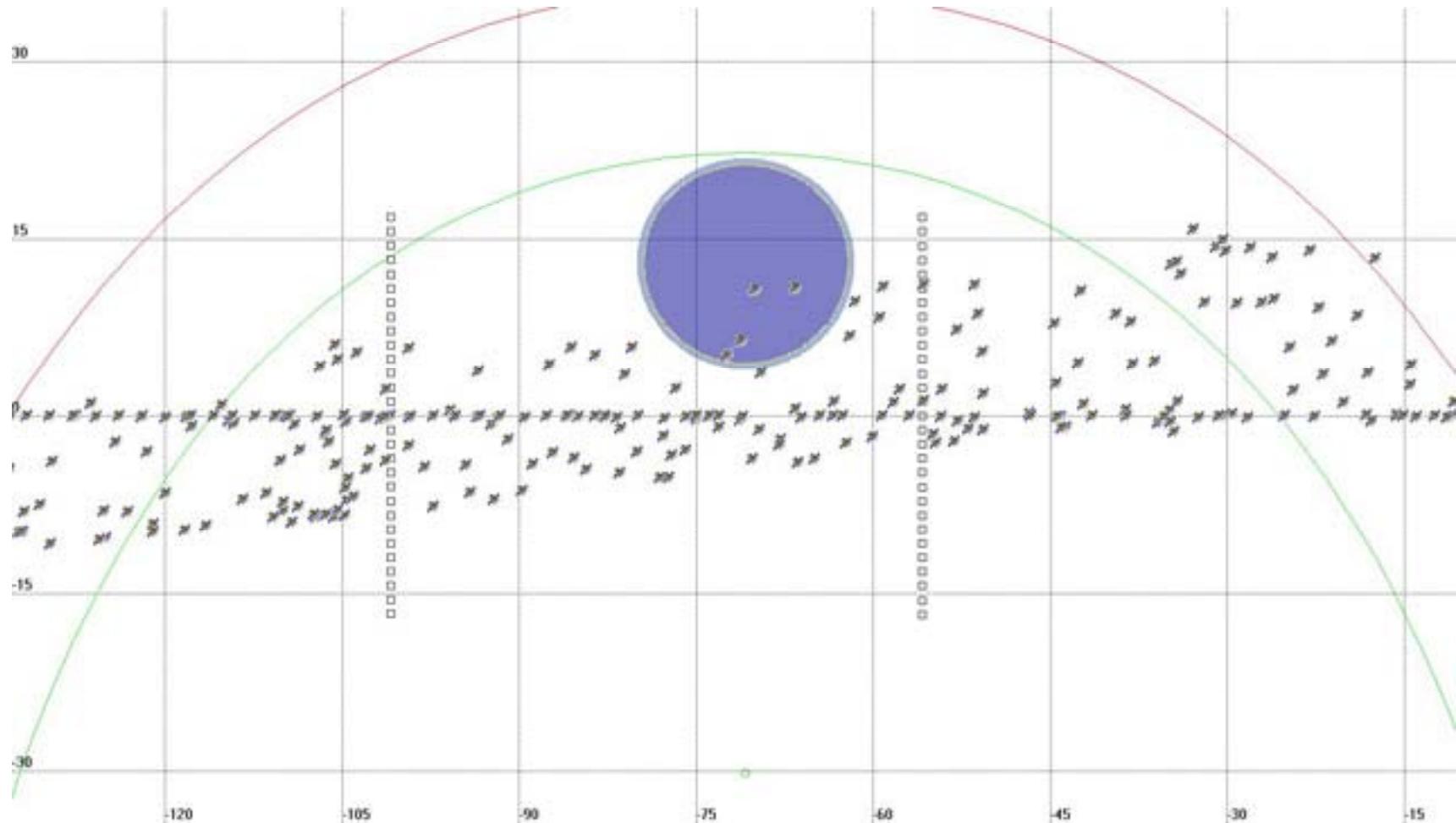


## Examples of Detections





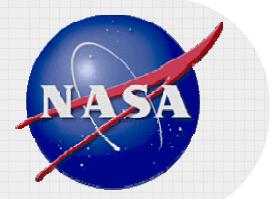
## Use of STK for Observation Planning





## GEO Debris Observations with Two Telescopes

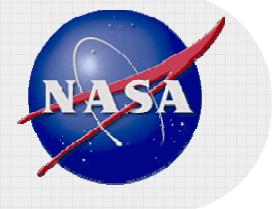
- Limitations of single telescope – incomplete survey or short arc for orbit.
- March 2007 – began observations with two telescopes at Cerro Tololo, Chile:
  - MODEST survey telescope: 0.6/0.9-m Schmidt
  - CTIO 0.9-m telescope for follow-up
- Goal – characterization of complete sample of faint GEO debris selected on basis of angular rates and brightness.
  - Orbits
    - What fraction of objects selected on basis of magnitude and angular rates are really at GEO?
  - Brightness variations
  - Colors in standard astronomical filters.



## CTIO 0.9-m

- **0.9-m Cassegrain**
- **0.22 deg FOV (small!)**
- **Track objects at their angular rates.**
- **Existing telescope and CCD.**
- **Debris project can obtain several weeks during the year of time on this telescope for GEO observations.**





## Techniques and Results

- Typical time between last MODEST observation and 1<sup>st</sup> 0.9-m observation less than 20 minutes. MODEST observations fit to circular orbit and prediction generated for 0.9-m.
- Recovery rate of MODEST detections on 0.9-m greater than 80%.
- After initial acquisition and follow-ups on 0.9-m: determine full six parameter orbit including eccentricity.
- Track objects from night to night
  - Longest track is six nights on multiple objects in March 2007, March & August 2008.
  - Primary reason for loss of track is object moves too far east or west to be tracked from Chile.

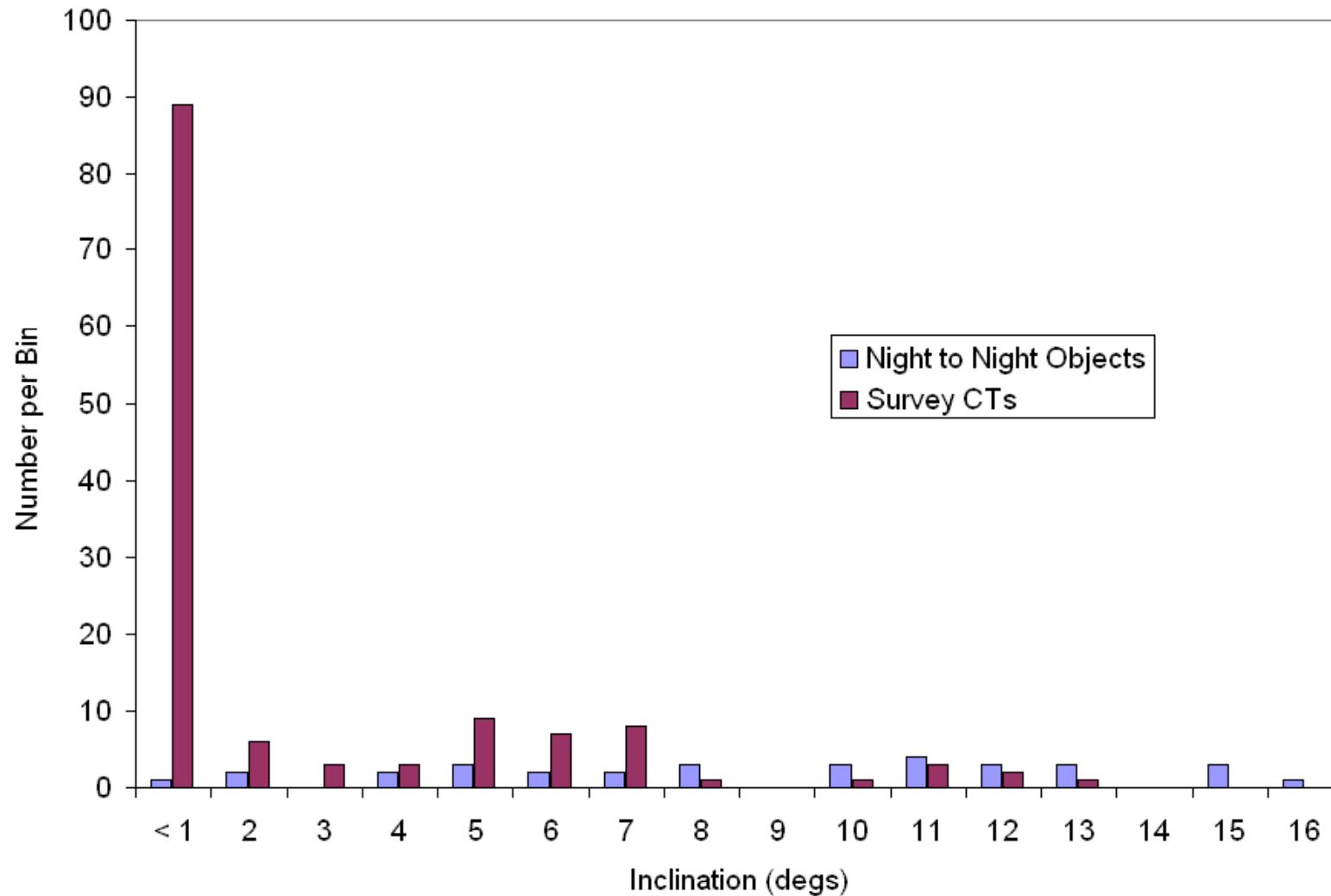


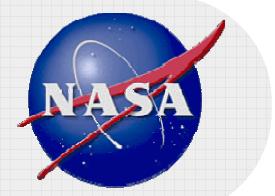
## Statistical Analysis

- **Use all available information from all sources to determine distribution of all objects found in MODEST survey strips.**
- **Bright objects found by MODEST not followed up by 0.9-m due to time constraints:**
  - Magnitudes from MODEST survey observations.
  - Orbits from public Space Command catalog.
- **Faint objects found by MODEST and followed up by 0.9-m:**
  - Magnitudes from MODEST survey observations to ensure on same system as bright objects (CTs) from catalog.
  - Orbits from full six parameter fit from both MODEST and 0.9-m observations.
- **A few bright objects with known orbits tracked by 0.9-m and orbits compared with catalog. Excellent agreement.**
- **Following slides show March 2007 data – analysis complete.**

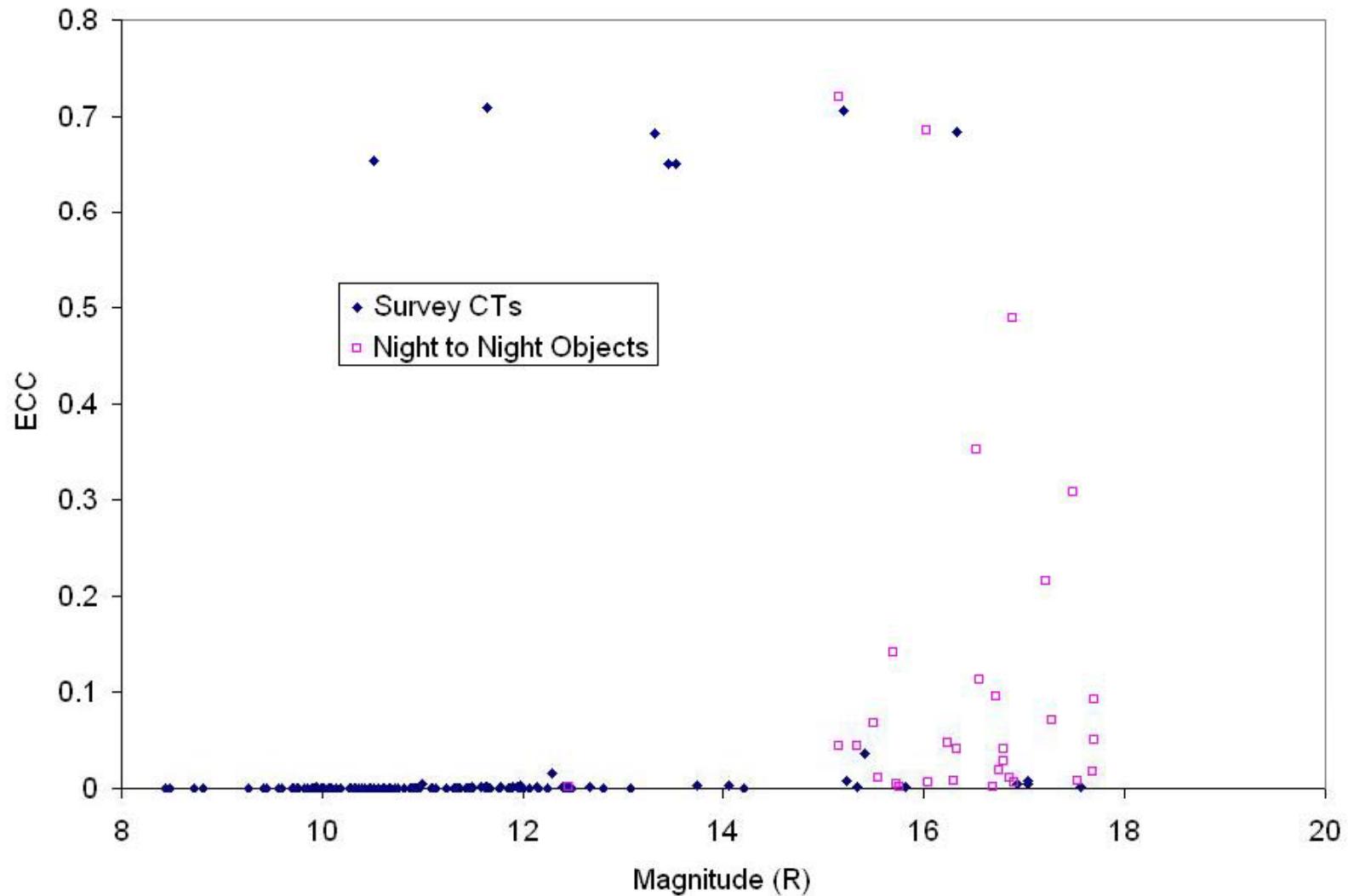


## Inclination Distribution



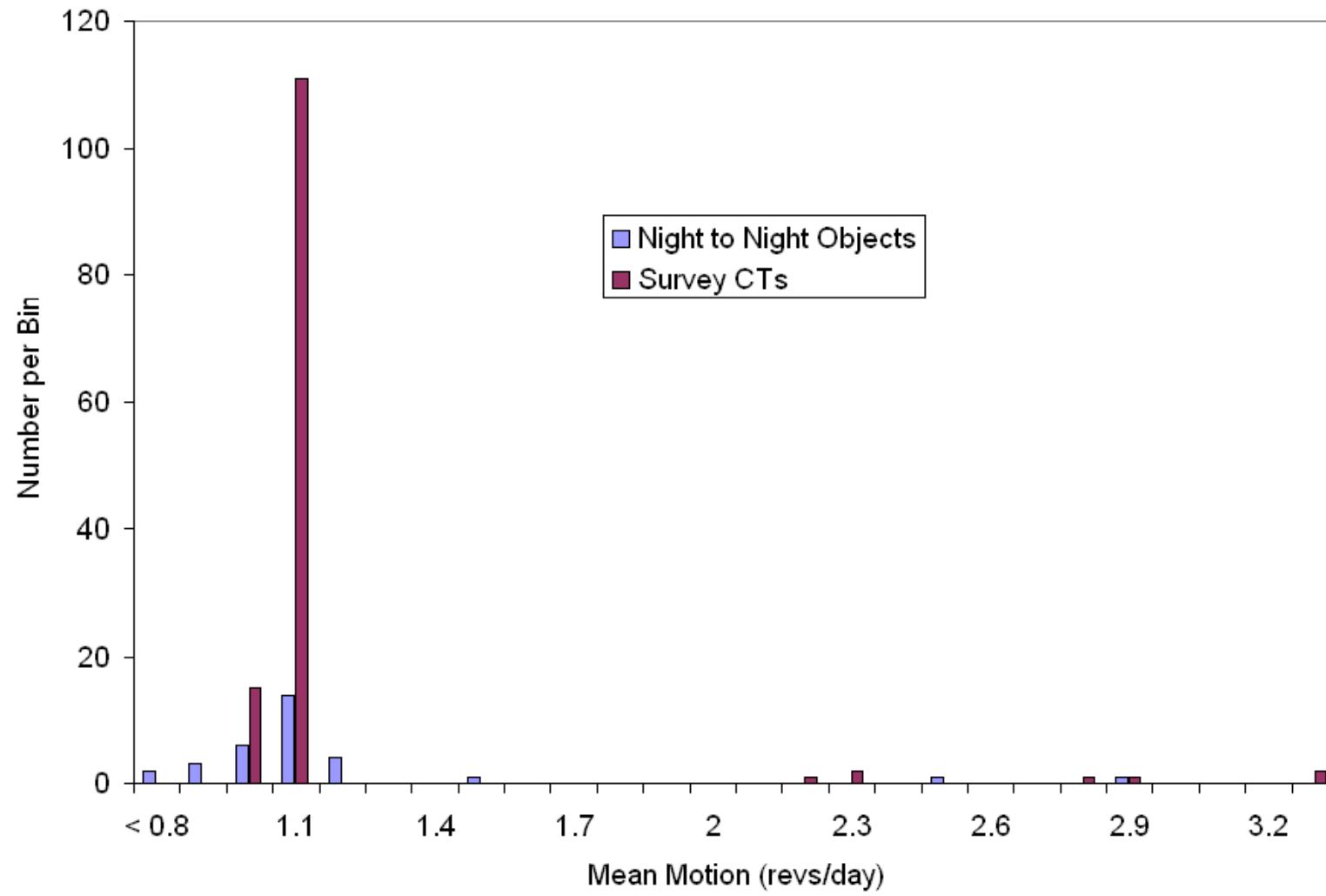


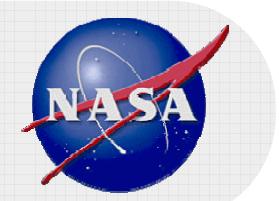
## Magnitude versus Eccentricity



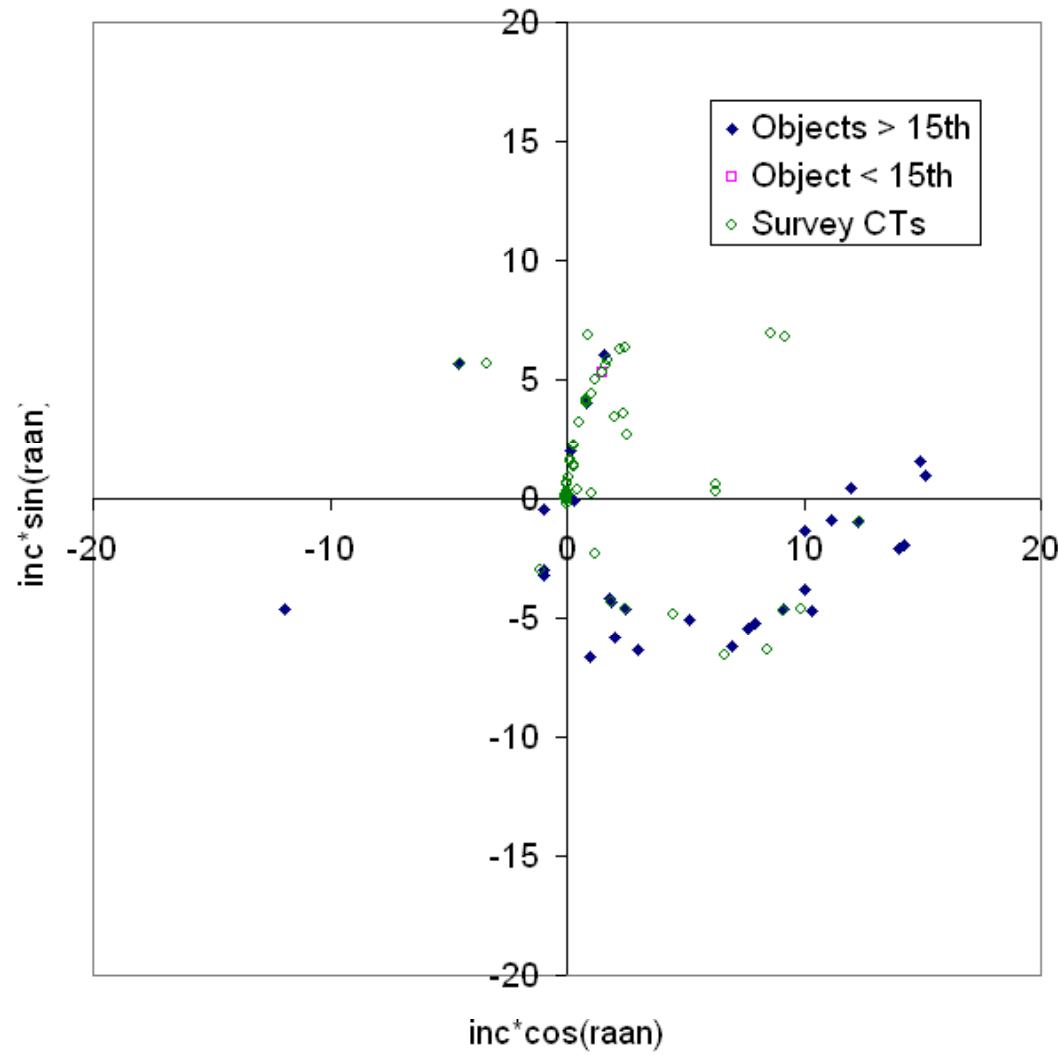


## Mean Motion Distribution



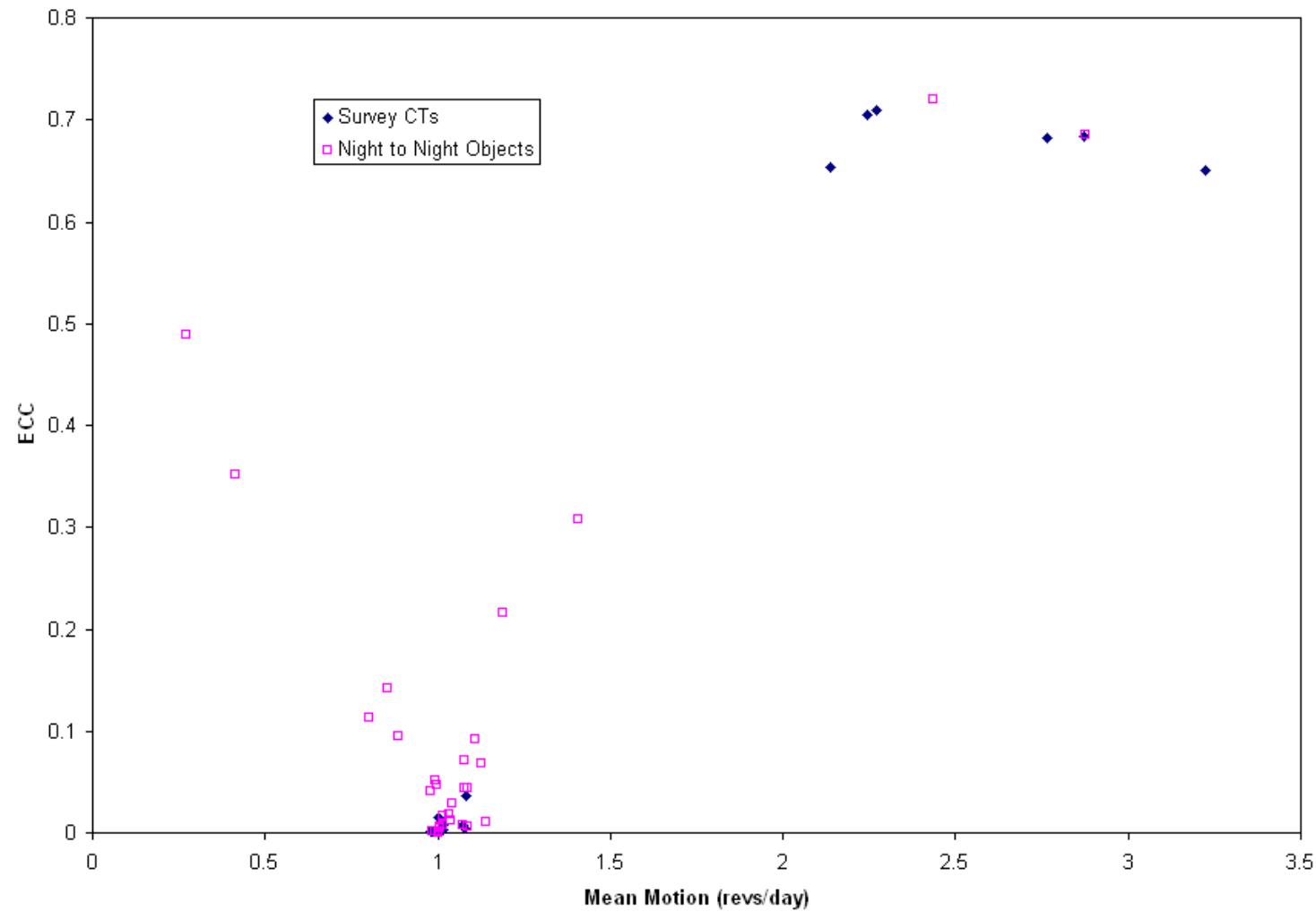


## Polar Plot





## Mean Motion vs Eccentricity





## Future

- **Most pressing goal - follow-up observations after several weeks with MODEST – determine A/M using ODTK.**
- **Goal – characterization of complete sample of faint GEO debris selected in a well defined, consistent manner.**